



VIIRS Nighttime Lights Development Update

Kimberly Baugh
Earth Observation Group (EOG)
CIRES - University of Colorado, USA
NOAA National Centers for Environmental Information (NCEI), USA
Kim.Baugh@noaa.gov

Chris Elvidge - NOAA NCEI, USA
Mikhail Zhizhin - CIRES - University of Colorado, USA
Feng Chi Hsu - CIRES - University of Colorado, USA
Tilottama Ghosh – CIRES – University of Colorado, USA

Nighttime Lights Composites

What are they?

A nighttime lights composite is made to serve as a baseline of persistent light sources.

Composites are made as an average of the highest quality nighttime lights imagery over desired time period – usually monthly or annually.

“Stable Lights” composites have ephemeral light sources and non-light (background) areas are removed from a composite.

EOG group is producing current monthly cloud-free/no-moon DNB nighttime lights composites and is doing algorithm development to turn these in to Stable Lights composites.

Nighttime Lights Composites

Processing Steps

Flag input DNB data so only the “highest quality” nighttime data gets averaged into a composite. Currently defined as:

- Cloud-free (using the VIIRS cloud-mask (VCM) product)
- Nighttime with solar zenith angles greater than 101
- Not affected by moonlight (lunar illuminance < 0.0005 lux)
- Middle of swath (DNB has increased noise at edge of scan)
- Free of lights from lightning
- Free of “lights” from South Atlantic Anomaly

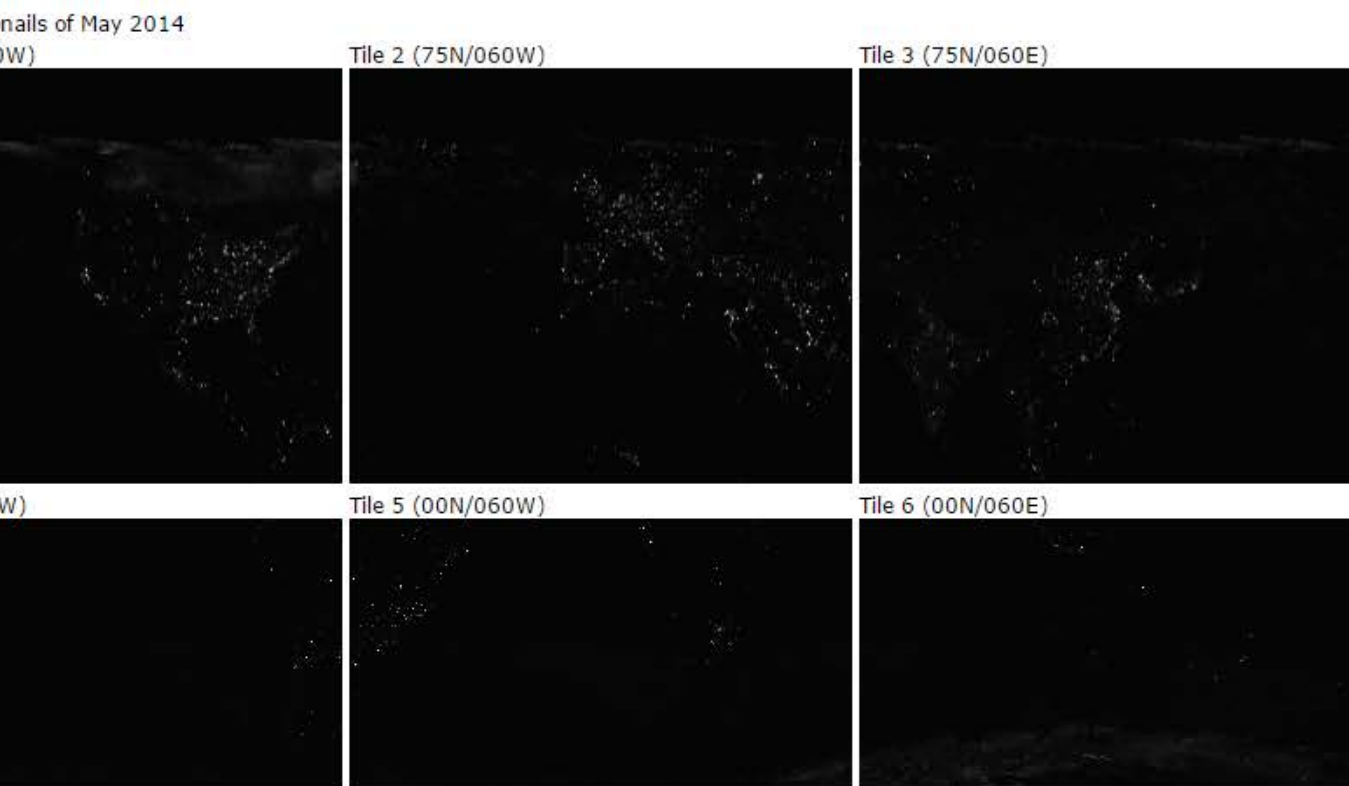
Create annual average DNB composite products and histograms of individual observations

Use annual histograms to remove DNB outliers (ephemeral lights and other sensor noise)

Identify and remove background (non-light) areas

Create TOA and atmospherically-corrected DNB composites

Nighttime Lights Composites (Monthly DNB Products)



- Monthly DNB nighttime lights composites are available online
- Globe is cut into 6 tiles to reduce individual file sizes
- These products still contain ephemeral lights and non-stationary lights (background).

Date: 11/30/2015/13:44:02

[Contract All](#)

5/October

5/September

5/August

5/July

5/June

http://www.ngdc.noaa.gov/eog/viirs/download_monthly.html

DNB Stray Light

Orthrup Grumman algorithm was implemented in the IDPS in August 2013.

Does a good job of mitigating stray light effects for visual interpretation.

Some issues for algorithm development within the stray light corrected region:

- Can under/over-correct, especially at transition into stray light and in Southern hemisphere

- Variance of data across scan is altered

- Correction quality is dependent on time from correction lookup table generation

- Stray light corrected regions are identified and processed separately

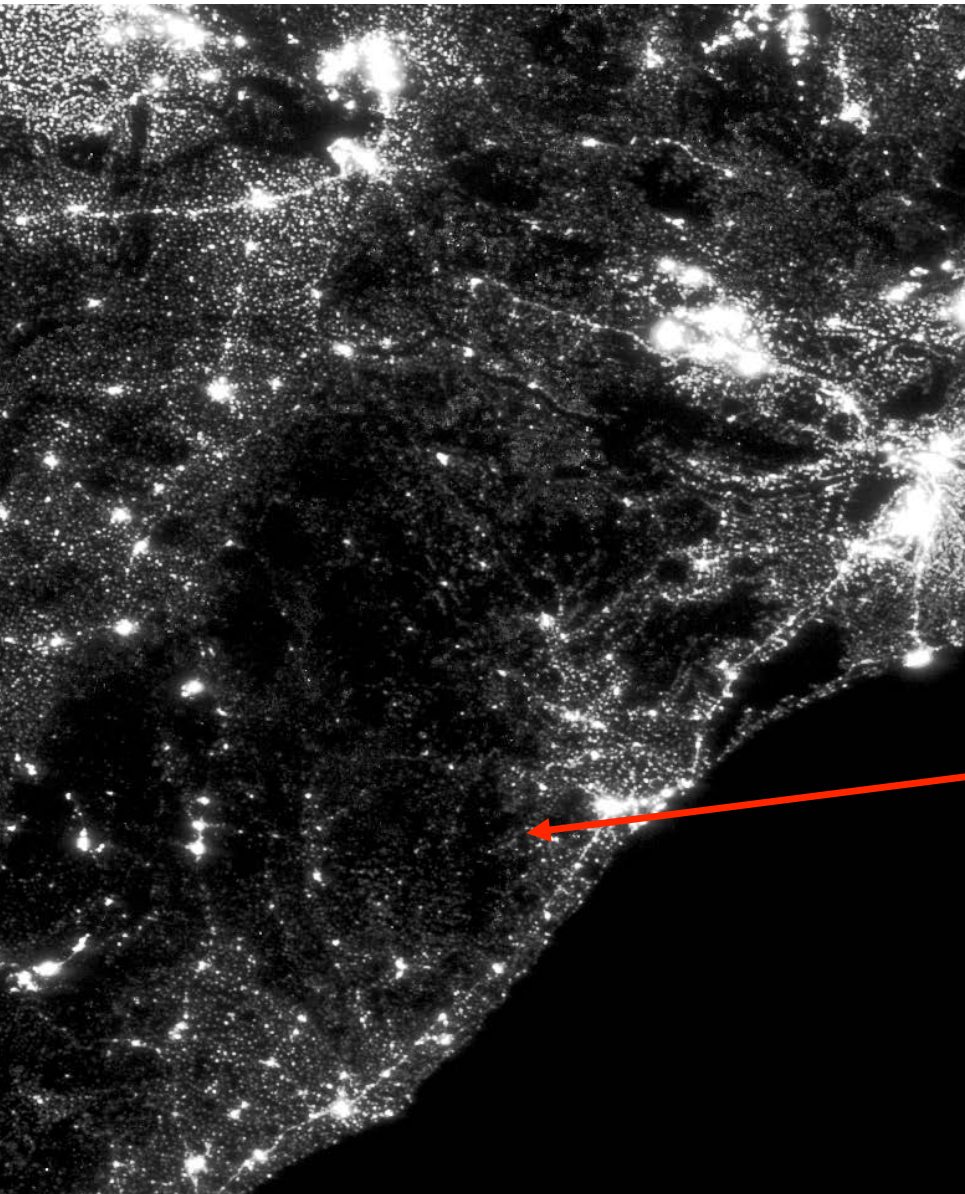


DNB Ephemeral Light Removal

Approach:

- Create histograms of DNB radiances using an extended time series (annual)
- Use histograms to identify and remove outliers
- Similar to algorithm developed for DMSP-OLS Stable Lights
- Advantages: This algorithm removes ANY outliers, including fires, boats, unfiltered-SAA, crosstalk, ...
- Disadvantages: Persistent flares and volcanic activity can remain. Method requires long time-series of data.

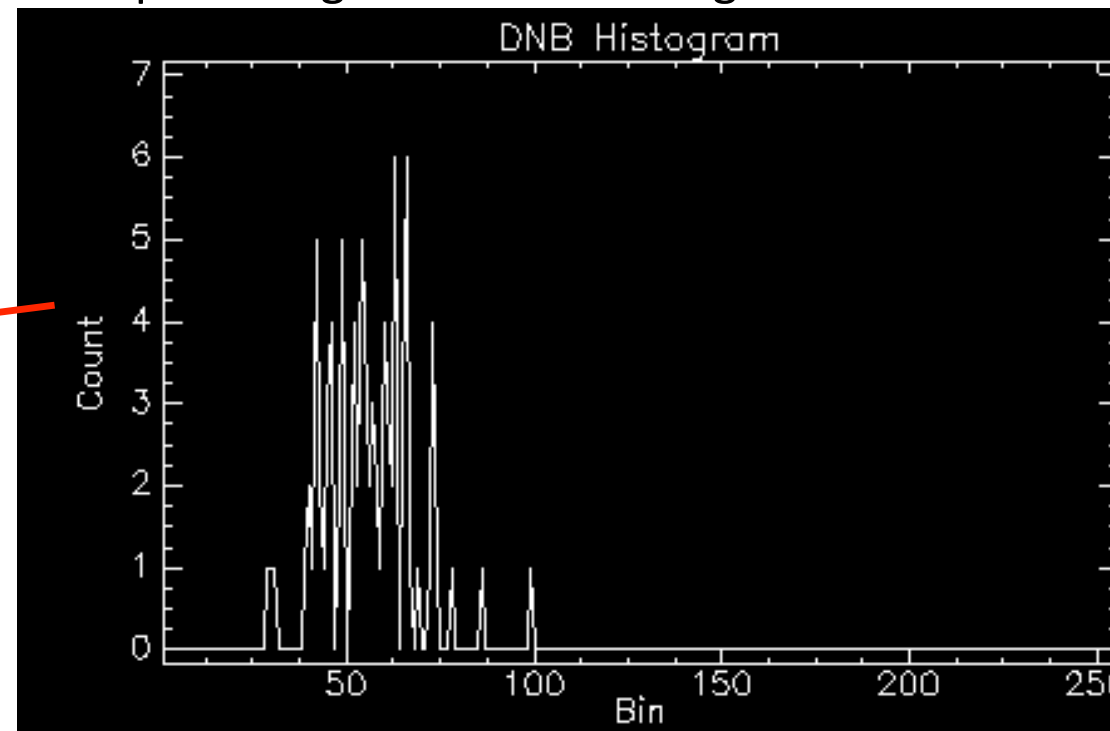
DNB Ephemeral Lights: Outlier Removal



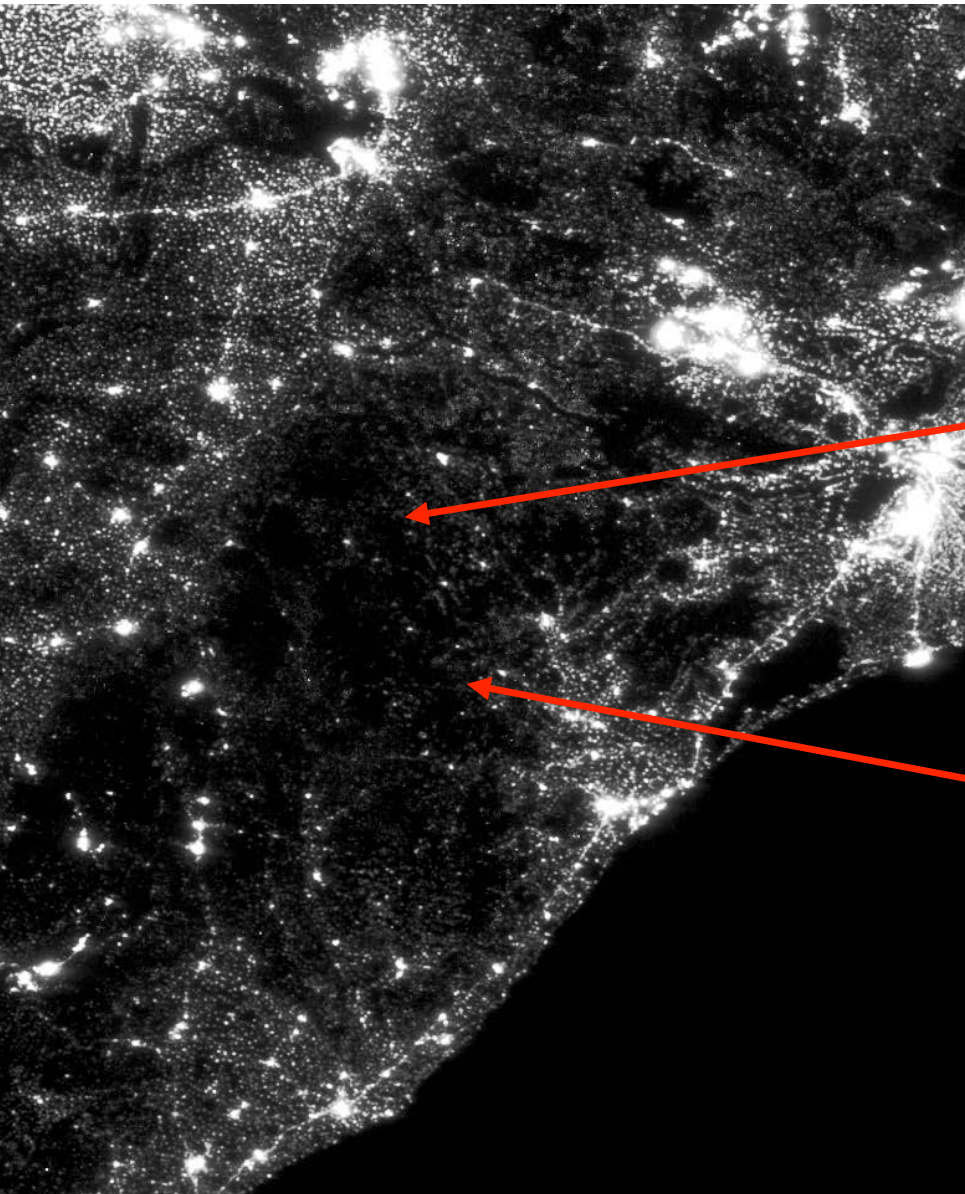
Odisha, India 2014 DNB Composite

- Histograms are made for each grid cell in composite
- DNB radiance values are placed in discrete bins on log transform. $\text{Bin} = \text{floor}(100 * (\log(1E9 * \text{Rad} + 1)))$

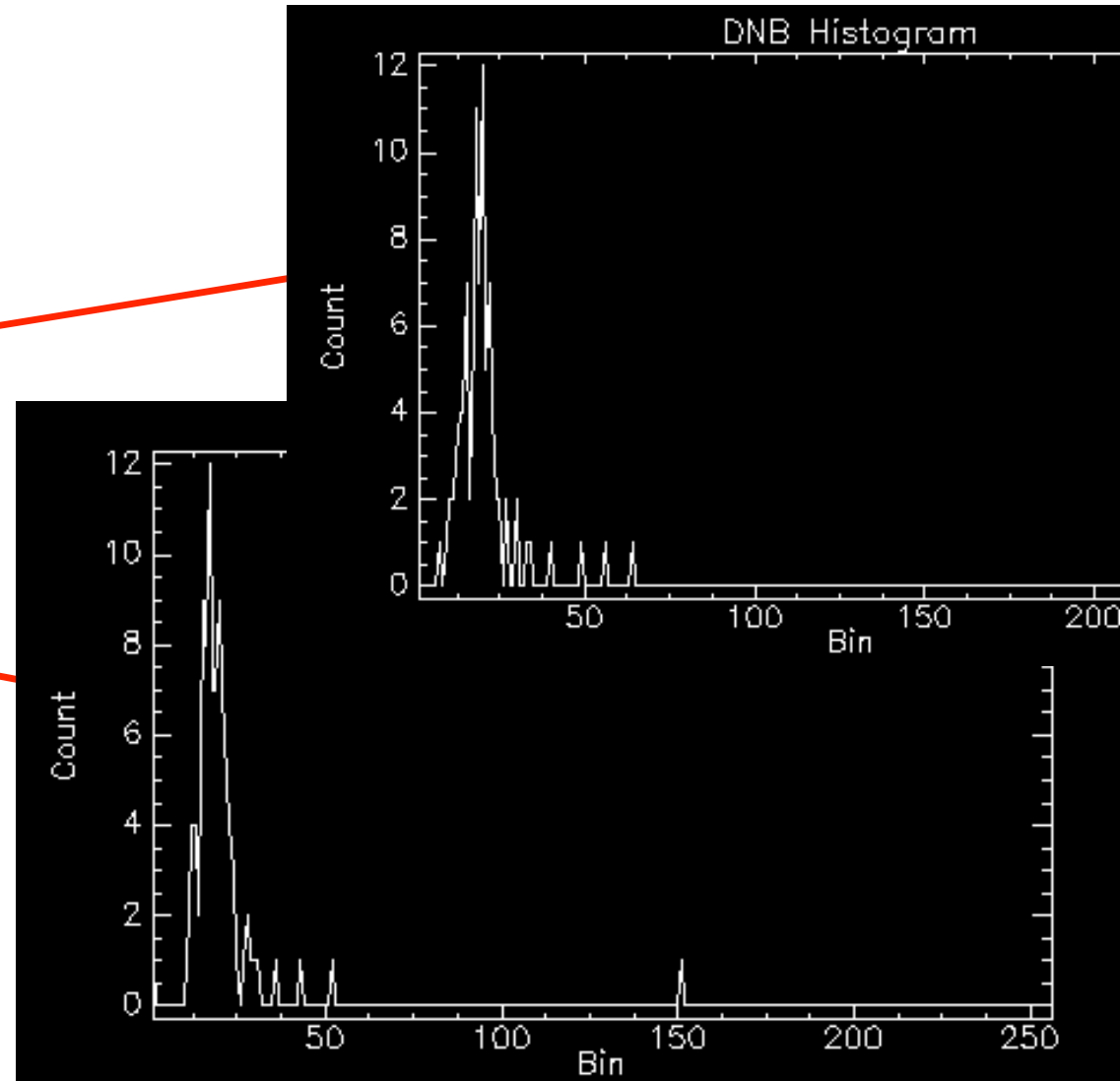
Example histogram of small village



DNB Ephemeral Lights: Outlier Removal



Example histograms of grid cells containing fires

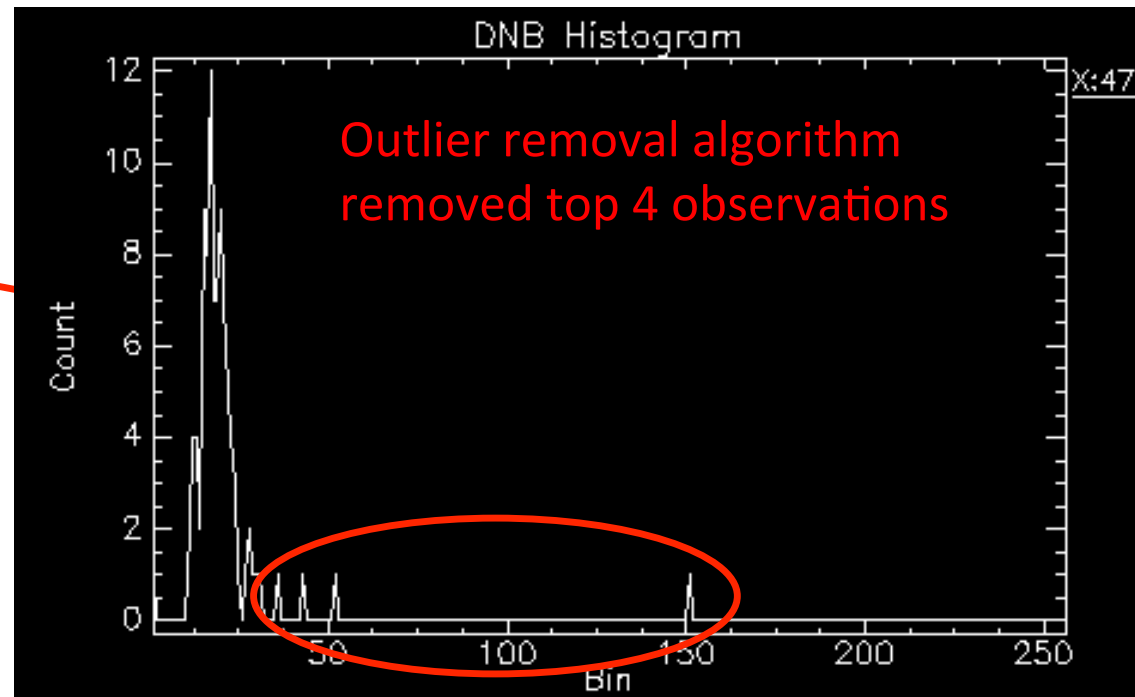


DNB Ephemeral Lights: Outlier Removal



Algorithm:

- 1) Compute standard deviation of observations
- 2) Remove highest observation
- 3) Re-compute standard deviation
- 4) Repeat steps 2-3 if difference in standard deviation threshold
- 5) Re-compute average of remaining observations



DNB Ephemeral Lights: Before Outlier Removal

with next slide

show regions with
activity return to
ground radiance levels
outlier removal



DNB Ephemeral Lights: After Outlier Removal

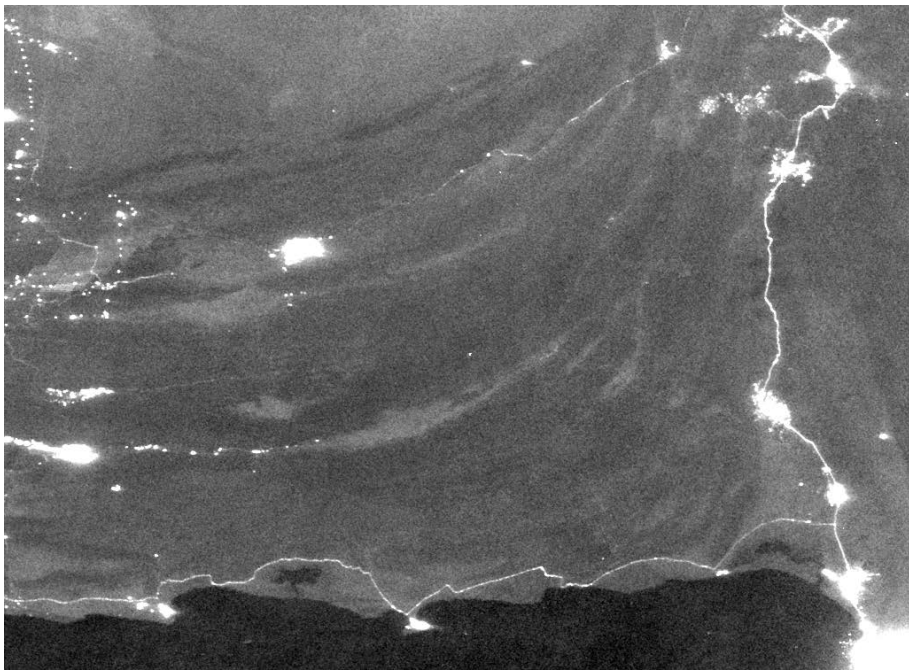
with previous slide

now regions with
ivity return to
und radiance levels
tlier removal



DNB Background Removal

- The DNB's detection limits are low enough, that even without moonlight present, nocturnal airglow can light up terrain and high albedo surfaces, making it challenging to separate dim lights from high albedo surfaces.

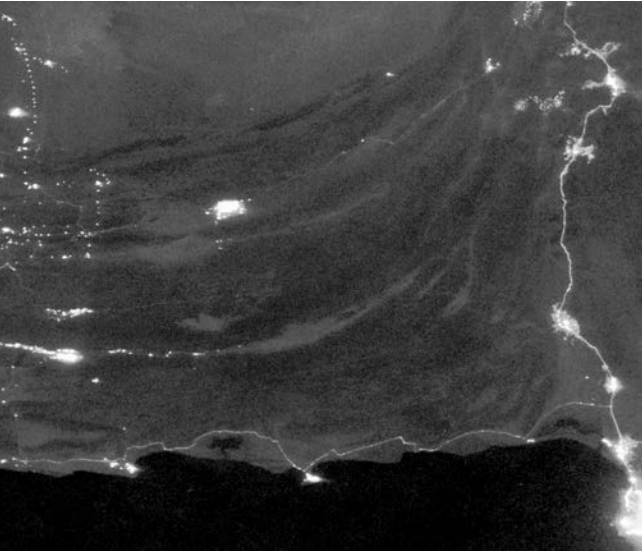


2014 DNB Composite over Southern Pakistan – some road features have lower average radiance values than no-light areas with high albedo

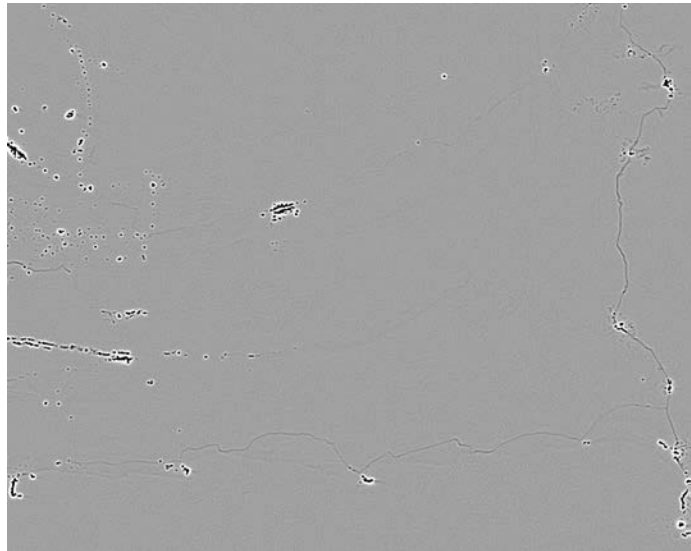


2014 DNB Composite over Himalayas – snow-covered peaks have higher average radiances than some of the villages

DNB Background Removal



2014 DNB Composite with
outliers removed



First derivative – areas close to
zero are background (gray)



2014 DNB Composite with
background masked using deriv
image

Radiance values of terrain surfaces can equal radiances of dim lights, but the values vary more slowly spatially than dim lights

First derivative, or gradient images of DNB composites lend well to thresholding to bring out nighttime lights

Initial testing shows most nighttime lights from cities/villages are retained, dim roads can get fragmented.

Atmospheric Correction for Nighttime DNB: Working with 6S

Inputs

TCO

- NOAA/OSPO TOAST

AOT

- NAAPS model (NPP VIIRS IVAOT)

TPW

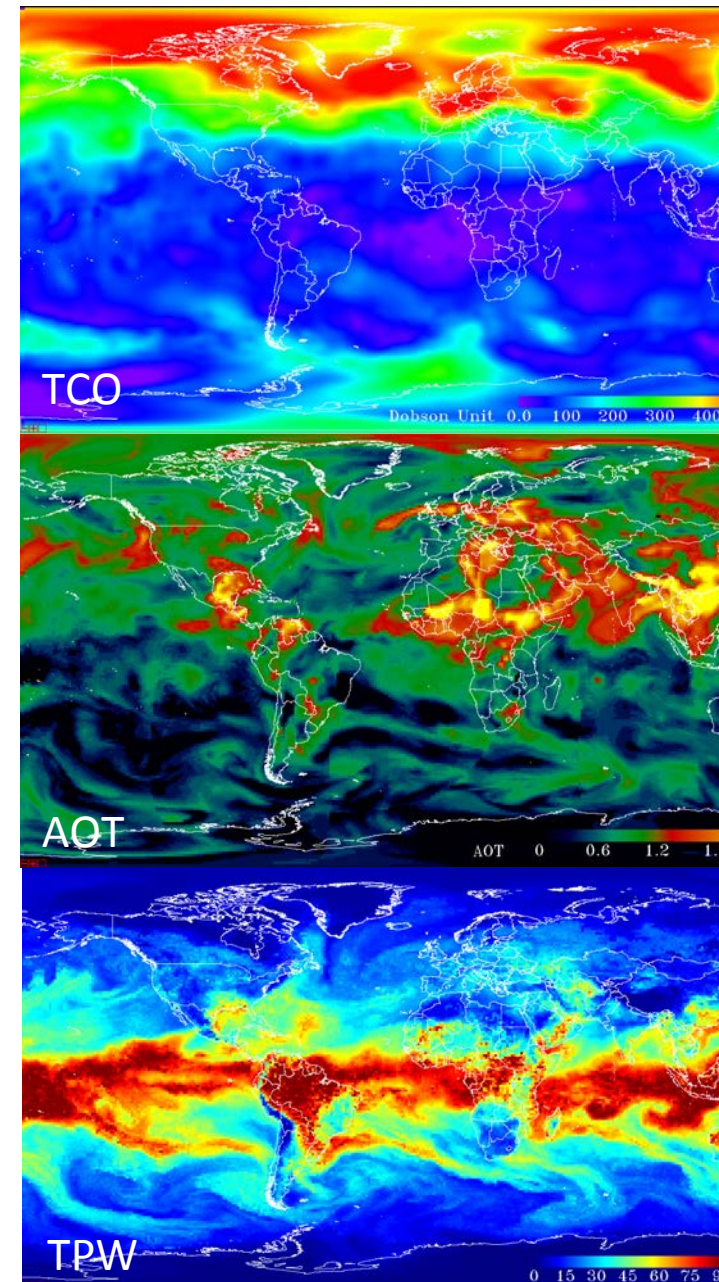
- NPP ATMS -> NOAA MIRS

Geometry

- SatZ, SatA, SolZ, SolA
- DEM: SRTM + GTOPO 30

Unified grid

- 1 degree Lat/Lon grid
- Confined by TOAST resolution / Save computation



Date represented: 2016/4/13

Radiative Transfer Model

$$\rho \uparrow^* = \rho \downarrow a + \rho \downarrow t / 1 - \rho \downarrow t S T(\theta \downarrow s) T(\theta \downarrow v)$$

For nocturnal self-emitting source under zero lunar illumination

- $\rho \downarrow a = 0$ (radiance in atmosphere)
- $\rho \downarrow t = 1$ (*reflectance of target*)
- $T(\theta \downarrow s) = 1$ (downwelling transmissivity)

Thus apparent radiance becomes

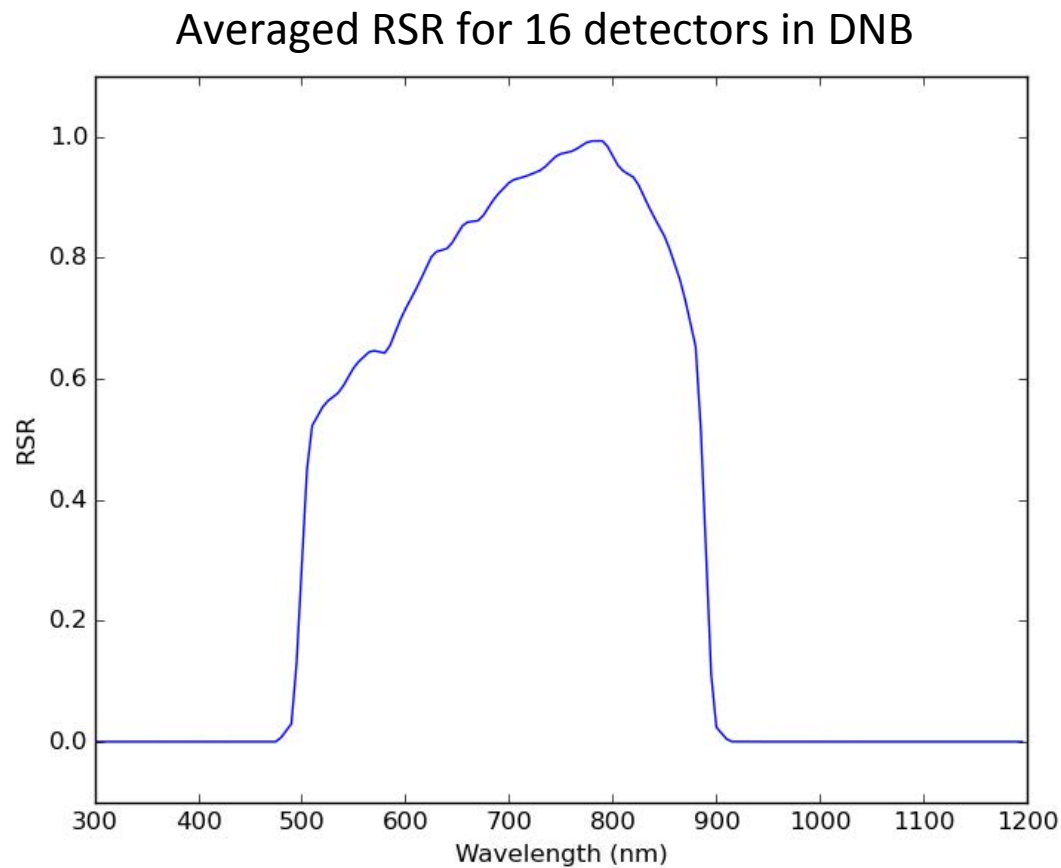
$$I \uparrow^* = I \rho \uparrow^* = I / 1 - S T(\theta \downarrow v)$$

Rewritten to isolate correction factor C

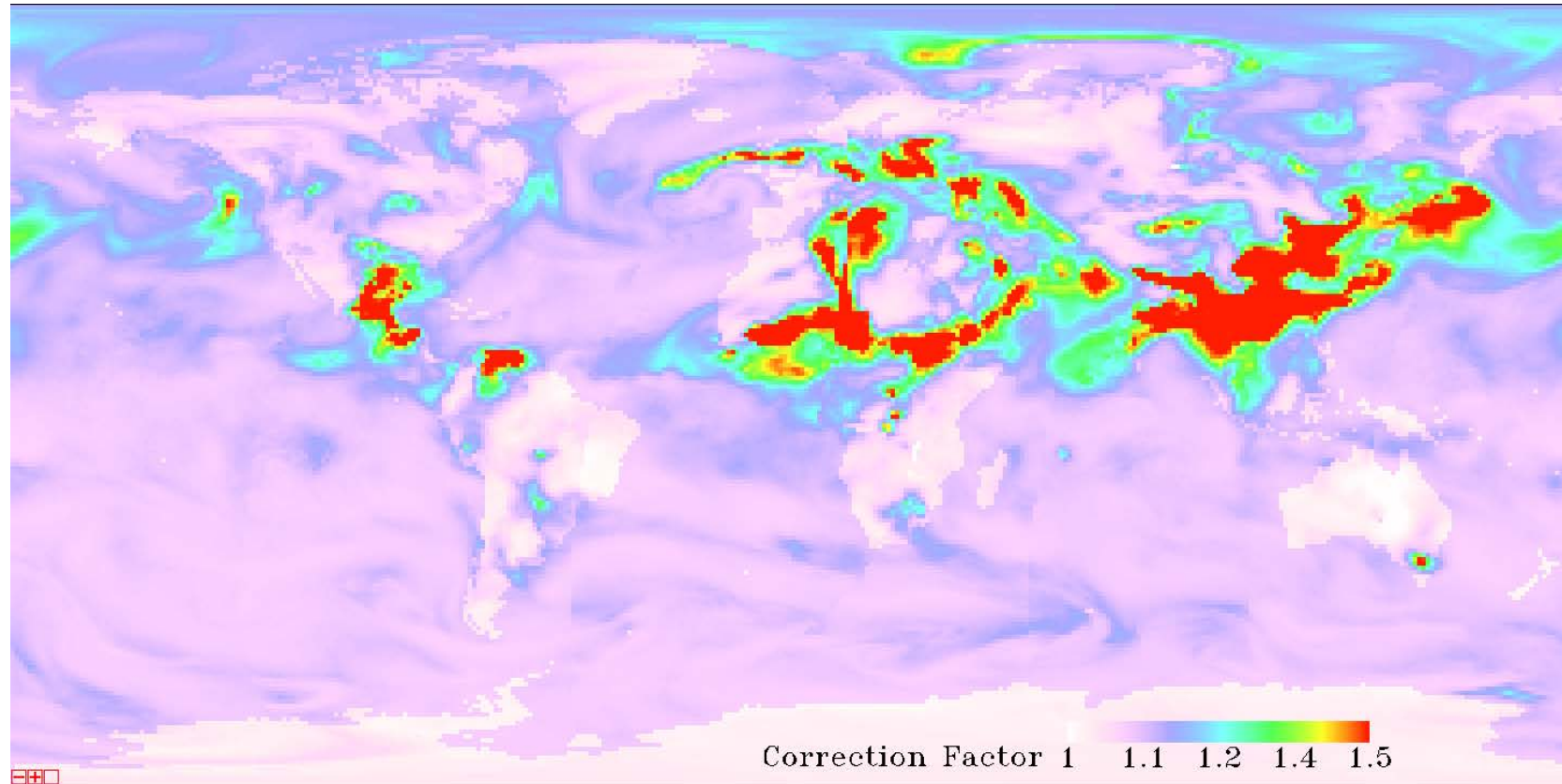
$$I = I \uparrow^* (1 - S) / T(\theta \downarrow v) = I \uparrow^* C$$

Band Averaging

- Consider the spectral sensitivity of DNB



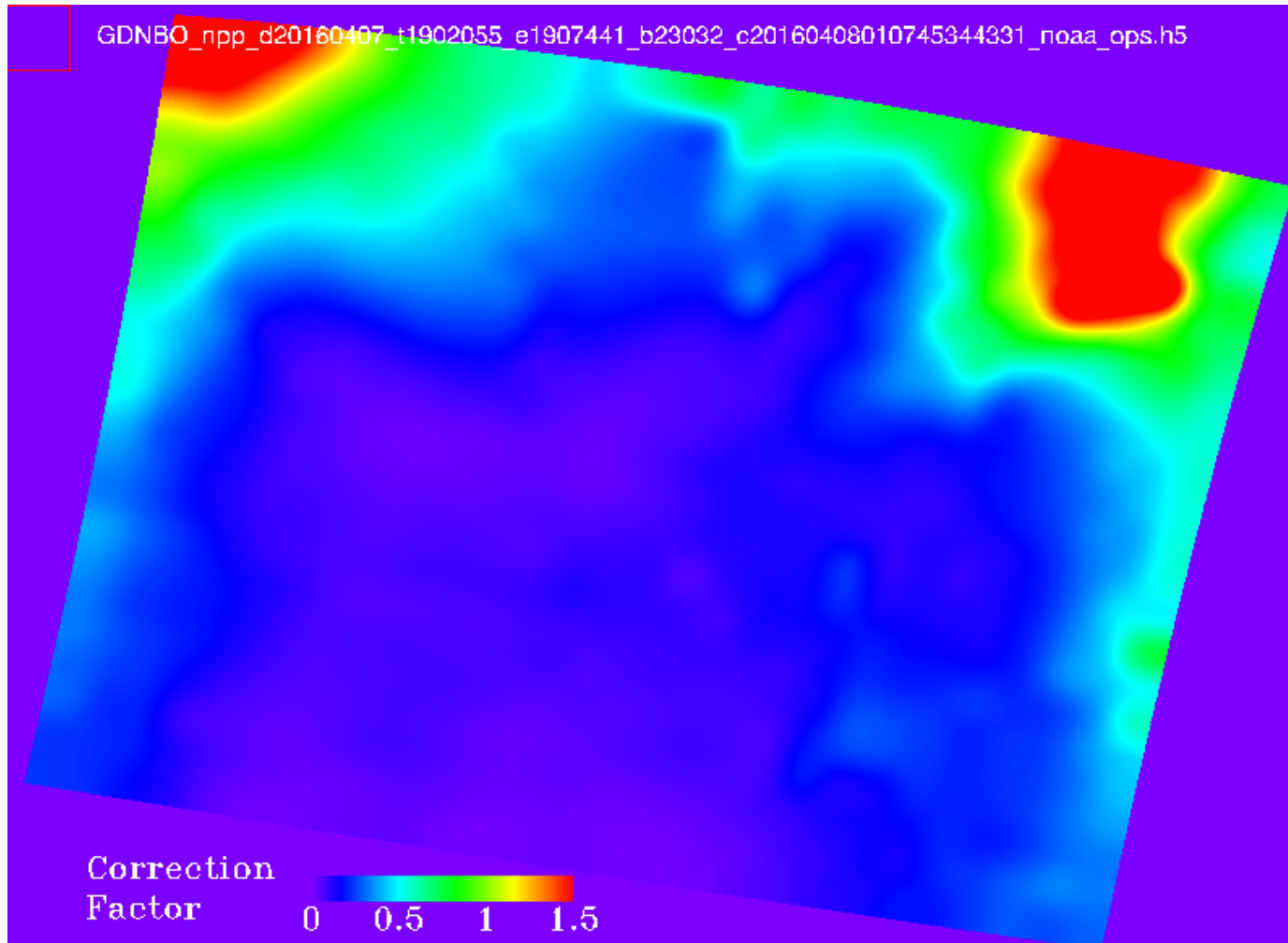
Global Correction Factor Grid



Date represented: 2016/4/13

This sample global grid is generated with fixed geometry properties
SATZ=0, SATA=0, SOLZ=150,SOLA=0

Aggregate Level Correction



Nighttime Lights Composites: Next Steps

- Finalize atmospheric correction algorithm
- Test outlier removal/background removal algorithms on aurora
- Add in Nightfire detections to identify locations of persistent flares and volcanos